

# **Stellar Rotation**

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## **Stellar Rotation:**

Stellar Rotation A. Slettebak, 2012-12-06 The International Astronomical Union Colloquium on Stellar Rotation was held at the Ohio State University in Columbus Ohio U S A from September 8th through 11 th 1969 Forty four scientists from Argentina Belgium Canada England Finland East and West Germany Italy Israel Japan The Netherlands and the United States attended and participated in the Colloquium The present volume which parallels the actual program closely contains the papers presented at the Colloquium plus most of the discussion following those papers The Colloquium was sponsored by the International Astronomical Union the Ohio State University and the National Science Foundation It is a pleasure to record my thanks to these organizations and especially to Dr Geoffrey Keller Dean of the College of Mathematics and Physical Sciences of the Ohio State University and to Prof C de Jager Assistant General Secretary of the International Astronomical Un ion for their kind cooperation I am also grateful to H A Abt J Hardorp R P Kraft Mrs A Massevitch M Plavec 1 W Roxburgh and E Schatzman of the Organizing Committee as well as A J Deutsch and G W Collins II for their help in planning the Colloquium all of them offered valuable suggestions toward organizing the program Stars Jean-Pierre Rozelot, 2008-11-27 The Sun and stars rotate in di erent ways and at di erent velocity rates The knowledge of how they rotate is important in understanding the formation and evolution of stars and their structure The closest star to our Earth the Sun is a good laboratory to study in detail the rotation of a G star and allows to test new ideas and develop new techniques to study stellar rotation More or less massive more or lessevolved objects however can have averydi erent rotation rate structure and history In recent years our understanding of the rotation of the Sun has greatly improved The Sun has a well known large scale rotation which can be m sured thanks to visible features across the solar disk such as sunspots or via spectroscopy In addition several studies cast light on di erential rotation in the convective zone and on meridional circulation in the radiative zone of the Sun Even the rotation of the core of the Sun can now be studied thanks to various methods such as dynamics of the gravitational moments and of course helioseismology through g modes analysis Moreover the magnetic eld is strongly linked to the matter motions in the solar plasma. The solar magnetic eld can be measured only at the surface or in theupperlayers Itistheproductoftheinternaldynamoorofthelocaldynamos if they exist in any case magnetic eld and rotation cannot thus be separated **Stellar Rotation** Arnold Hanslmeier, Roman Brajša, 2024-06-25 This textbook highlights mechanisms properties observations and research progress of solar and stellar rotation The rotation of stars is a basic parameter that enters into the dynamo mechanisms that generates stellar activity. The faster stars rotate the higher stellar activity Moreover the rotation can also be used to determine stellar ages Only in the case of the Sun we can directly observe details and therefore determine e g the differential rotation Thus the Sun is a prototype for studying other stars The textbook presents techniques that enable to determine rotation on stars and give a comparison of how stars rotate according to their spectral type and evolution Special cases are given such as rapidly rotating stars that are strongly deformed into

ellipsoids The textbook is intended for professionals and students in Astrophysicists as well as readers interested in science Physics, Formation and Evolution of Rotating Stars Andre Maeder, 2008-12-19 Rotation is ubiquitous at each in general step of stellar evolution from star formation to the final stages and it affects the course of evolution the timescales and nucleosynthesis Stellar rotation is also an essential prerequisite for the occurrence of Gamma Ray Bursts In this book the author thoroughly examines the basic mechanical and thermal effects of rotation their influence on mass loss by stellar winds the effects of differential rotation and its associated instabilities the relation with magnetic fields and the evolution of the internal and surface rotation Further he discusses the numerous observational signatures of rotational effects obtained from spectroscopy and interferometric observations as well as from chemical abundance determinations helioseismology and asteroseismology etc On an introductory level this book presents in a didactical way the basic concepts of stellar structure and evolution in track 1 chapters The other more specialized chapters form an advanced course on the graduate level and will further serve as a valuable reference work for professional astrophysicists **Stellar Rotation** Jean-Louis Tassoul, 2000-04-13 Like the Earth and planets stars rotate Understanding how stars rotate is central to modelling their structure formation and evolution and how they interact with their environment and companion stars This authoritative volume first published in 2000 provides a lucid introduction to stellar rotation and the definitive reference to the subject It combines theory and observation in a comprehensive survey of how the rotation of stars affects the structure and evolution of the Sun single stars and close binaries This book will be of primary interest to graduate students and researchers studying solar and stellar rotation and close binary systems It will also appeal to those with a more general interest in solar and stellar physics star formation binary stars and the hydrodynamics of rotating fluids including geophysicists planetary scientists and plasma physicists NASA Thesaurus, 1998 Contains the authorized subject terms by which the documents in the NASA STI Database are indexed and retrieved Old Stellar Populations Santi Cassisi, Maurizio Salaris, 2013-08-29 The book discusses the theoretical path to decoding the information gathered from observations of old stellar systems It focuses on old stellar systems because these are the fossil record of galaxy formation and provide invaluable information on the evolution of cosmic structures and the universe as a whole The aim is to present results obtained in the past few years for theoretical developments in low mass star research and in advances in our knowledge of the evolution of old stellar systems A particularly representative case is the recent discovery of multiple stellar populations in galactic globular clusters that represents one of the hottest topics in stellar and galactic astrophysics and is discussed in detail Santi Cassisi has authored about 270 scientific papers 150 of them in peer reviewed journals and the title Evolution of Stars and Stellar Populations

<u>Fundamentals of Astrophysics</u> Mr. Rohit Manglik,2024-09-07 EduGorilla Publication is a trusted name in the education sector committed to empowering learners with high quality study materials and resources Specializing in competitive exams and academic support EduGorilla provides comprehensive and well structured content tailored to meet the needs of students

across various streams and levels Theoretical and Observational Consequences of Stellar Rotation and Magnetic Fields in From the sun's energy source to the formation of the solar system Dan Stellar Winds Richard Ignace, 1996 Bar-Zohar, 2006-12-04 The latest observation of hundreds of exoplanets and the discovery of supermassive black hole at the center of many galaxies set the foundation for the theory presented in this book. The theory suggest that the sun and stars energy source is not from fusion but instead from magnetic fields spreads in the galaxy by the supermassive black hole at the center of every galaxy This idea changes every aspect of astronomy and cosmology The big bang is no longer necessary to explain the source of the mass in the universe and the expansion of the universe According to this theory the matter in the universe is created in the cores of stars by conversion of energy to mass The expansion of the universe is induced by the rapid formation of new galaxies Stars grow slowly and gradually over tens of billion of years by conversion of energy to mass The gradual growth of stars and the planet search programs that found hundreds of nearby planets indicate that stars are born from planets This invalidates the solar nebula hypothesis as the source of the stars and the solar system Stars fluctuate from a main sequence state to a red giant state They stay in the main sequence when they receive strong magnetic fields and they turn into a red giant when the magnetic fields are weakened The sun also fluctuated from a main sequence to a red giant When the sun was a red giant it had strong solar wind that supplied the material to created the planets The solar system contains hard evidence that the sun was a red giant those are short lived isotopes and chondrules. The fact that there is hard evidenced to a red giant sun confirm this theory Highlights of this theory include the following 1 The sun energy source is from magnetic fields from the galactic center 2. The heat induced by the magnetic fields leads to high energy collision between particle in the sun core that creates new particle and increase the sun mass 3 All the stars in the galaxy create new mass so the total mass and the size of the galaxy is increasing 4. The stars in the galaxy eject dust that freefall to the galactic center supermassive black hole Thorough the dynamo effect the gravitational potential energy of the debris and dust is converted to magnetic fields 5 As the galaxy mass and size increase globular clusters are detached form the main galaxy to create new galaxies 6 Galaxies spawn new galaxies and the total number of galaxies in the universe increase 7 The universe expands and accelerates from the increase in the number of the galaxies 8 The Big Bang cosmological model is replaced by a new cosmological model that resembles the steady state theory 9 Stars grow gradually from conversion of energy to mass 10 Stars are born from planets they first grow by accretion and then by conversion of energy to mass 11 Stars fluctuate from main sequence to a red giant When the magnetic fields are strong the star is in the main sequence when the magnetic fields are weakened the star turn to a red giant 12 The sun was a red giant 4 6 billions years ago 13 The planets were created from the strong solar wind of the red giant sun **Solar Surface Magnetism** R. J. Rutten, C.J. Schrijver, 2012-12-06 Observations of the solar magnetic field are largely confined to the radiation emitted from the photosphere the thin layer of the solar atmosphere which we call the solar surface It is from solar surface observations that

we must infer the internal structure and the internal magnetohydrodynamic processes that lead to the multitude of fascinat ing phenomena of solar magnetic activity and from solar surface observations we must also infer the interplay of convection and magnetism that regulates field dispersal drives the heating of the outer atmospheric plasma and generates the solar wind There is much to be learned from solar surface magnetism in physics and astrophysics currently there are rapid developments in this exciting field The workshop of which this volume contains the proceedings aimed at a synthesis between observers and theorists both with regard to the discrete elements that are the building blocks of solar magnetism and with regard to the larger scale spatial and temporal patterns in which the magnetic elements emerge and disappear The workshop was held during November 1 5 1993 in Soesterberg The Netherlands The fifty participants took a very active part in making the workshop quite a lively one The articles in these proceedings cover most of the oral and poster presentations excepting a dozen soon to be published elsewhere Astronomy And Relativistic Astrophysics: New Phenomena And New States Of Matter In The Universe - Proceedings Of The Third Workshop (Iwara07) Horst Stocker, Cesar Augusto Zen Vasconcellos, Valdir B Bezerra, Marcelo J Reboucas, Walter Greiner, Bardo E J Bodmann, 2010-04-15 This book presents scientific research on the central theme of new states of matter and new phenomena in the universe The topics covered range from the big bang through topics including the formation of exotic stars black holes and the plasma of guarks and gluons by heavy ion reactions to the influence of dark matter and dark energy in the evolution of the universe Scientific interest in these themes has been growing together with the development of major projects such as AUGER LHC FERMI Telescope FAIR GSI and GEO LIGO the diversity and scope of research on such themes has been on the rise The content is detailed enough to capture the interest of experts in the field and is useful for future explorations on these fascinating Star-Disk Interaction in Young Stars (IAU S243) International Astronomical Union. Symposium, 2007 themes Reviews latest star disk interaction region in young stars for graduate students and researchers Heliophysics: Active Stars, their Astrospheres, and Impacts on Planetary Environments Carolus J. Schrijver, Frances Bagenal, Jan J. Sojka,2016-03-17 Heliophysics is a fast developing scientific discipline that integrates studies of the Sun's variability the surrounding heliosphere and the environment and climate of planets This volume the fourth in the Heliophysics collection explores what makes the conditions on Earth just right to sustain life by comparing Earth to other solar system planets by comparing solar magnetic activity to that of other stars and by looking at the properties of evolving exoplanet systems By taking an interdisciplinary approach and using comparative heliophysics the authors illustrate how we can learn about our local cosmos by looking beyond it and in doing so also enable the converse Supplementary online resources are provided including lecture presentations problem sets and exercise labs making this ideal as a textbook for advanced undergraduate and graduate level courses as well as a foundational reference for researchers in the many subdisciplines of helio and astrophysics Solar-Type Activity in Main-Sequence Stars Roald E. Gershberg, 2005-07-12 The first comprehensive

monograph on this active and productive field of research investigates solar type activity amongst the large spectrum of low and middle mass main sequence stars and presents the subject in a systematic and comprehensive fashion Galaxies Tommy Wiklind, Bahram Mobasher, Volker Bromm, 2012-12-15 New observations of the period between the cosmic recombination and the end of reionization are posing intriguing questions about where the first generations of stars were formed how the first galaxies were assembled whether these galaxies have low redshift counterparts and what role the early galaxies played in the reionization process Combining the new observational data with theoretical models can shed new light on open issues regarding the star formation process its role in the reionization of the Universe and the metal enrichment in galaxies at those early epochs This volume brings together leading experts in the field to discuss our current level of understanding and what may come in the near future as our observational as well as theoretical tools improve The book confronts the theory of how the first stars black holes and galaxies formed with current and planned observations This synthesis is very timely just ahead of the establishment of major new facilities such as the James Webb Space Telescope JWST a next generation millimeter sub millimeter observatory in the Atacama desert ALMA and ground based Extremely Large Telescopes ELT Together they will revolutionize the study of the most distant objects in the Universe This volume is aimed at beginning graduate students but can also serve as a reference work for active researchers in the field Apart from presenting the fundamental concepts involved it also provides an introduction to the methods and techniques used The book will also be useful to anyone with an astrophysical background who needs an effective starting point for learning about the first stars and galaxies Rotation and Mixing in Stellar Interiors Marie-Jo Goupil, Jean-Paul Zahn, 1990 Analysis of Starlight J. B. Hearnshaw, 1990-04-19 This book presents a detailed pedagogical account of the equation of state and its applications in several important and fast growing topics in theoretical physics chemistry and engineering This book is the story of the analysis of starlight by astronomical spectroscopy. It describes the development of the subject from the time of Joseph Fraunhofer who in 1814 used a telescope mounted prism to observe the spectral light emitted from several bright stars He discovered that light was missing at certain colours wavelengths in the starlight and these so called spectral lines were subsequently shown to hold clues to the nature of the stars themselves. The book explains how the classification of stars using their line spectra developed into a major branch of astronomy whilst new methods in astrophysics made possible the approximate quantitative analysis of spectral lines in the 1920s and 1930s After the Second World War these techniques were considerably improved when computers were programmed to model the structure of the outer layers of stars Basic concepts in spectroscopy and spectral analysis are also covered and finally Dr Hearnshaw comments on the stellar spectroscopy of some individual star Literature 1997, Part 1 Astronomisches Rechen-InstitutARI,2013-11-11 Astronomy and Astrophysics Abstracts is devoted to the recording summarizing and indexing of astronomical publications throughout the world Two volumes are scheduled to appear per year Volume 67 records 10 903 papers covering besides the classical

fields of astronomy and astrophysics such matters as space flights related to astronomy lunar and planetary probes and satellites meteorites and interplanetary matter X rays and cosmic rays quasars and pulsars The abstracts are classified under more than one hundred subject categories thus permitting quick surveying of the bulk of material published on the same topic within six months For instance this volume records 119 papers on minor planets 155 papers on supernovae and 554 papers on cosmology 

Determination of Atmospheric Parameters of B-, A-, F- and G-Type Stars Ewa Niemczura, Barry Smalley, Wojtek Pych, 2014-07-10 This book introduces the theory of stellar atmospheres Almost everything we know about stars is by analysis of the radiation coming from their atmospheres Several aspects of astrophysics require accurate atmospheric parameters and abundances Spectroscopy is one of the most powerful tools at an astronomer s disposal allowing the determination of the fundamental parameters of stars surface temperature gravity chemical composition magnetic field rotation and turbulence These can be supplemented by distance measurements or pulsation parameters providing information about stellar interior and stellar evolution otherwise unavailable The volume is based on lectures presented at the Wroc aw s Spectroscopic School aimed at training young researchers in performing quantitative spectral analysis of low mid and high resolution spectra of B A and F type stars

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