# Supernovae –Type II

Massive stars at the end of their lifetimes develop iron cores. The core is degenerate and grows as shell burning continues. When the mass of the core exceeds the Chandrasekhar limit, the core starts to contract rapidly.

Two types of instabilities develop. First, electrons are captured by heavy nuclei, which reduces the presssure. Second, the degenerate matter is less sensitive to temperature changes, and the temperature rises quickly. Eventually, this leads to photodisintegration of iron nuclei. This reaction is endothermic, and absorbs about 2 MeV per nucleon. Both of these processes reduce the energy and the pressure in the core, which now collapses in almost free fall.

As the temperature rises even further, highly energetic photons break the helium nuclei into protons and neutrons, a reaction which absorbs about 6 MeV per nucleon.



# **Supernovae Physics Astronomical Monograph**

**M Woodhall** 

# **Supernovae Physics Astronomical Monograph:**

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