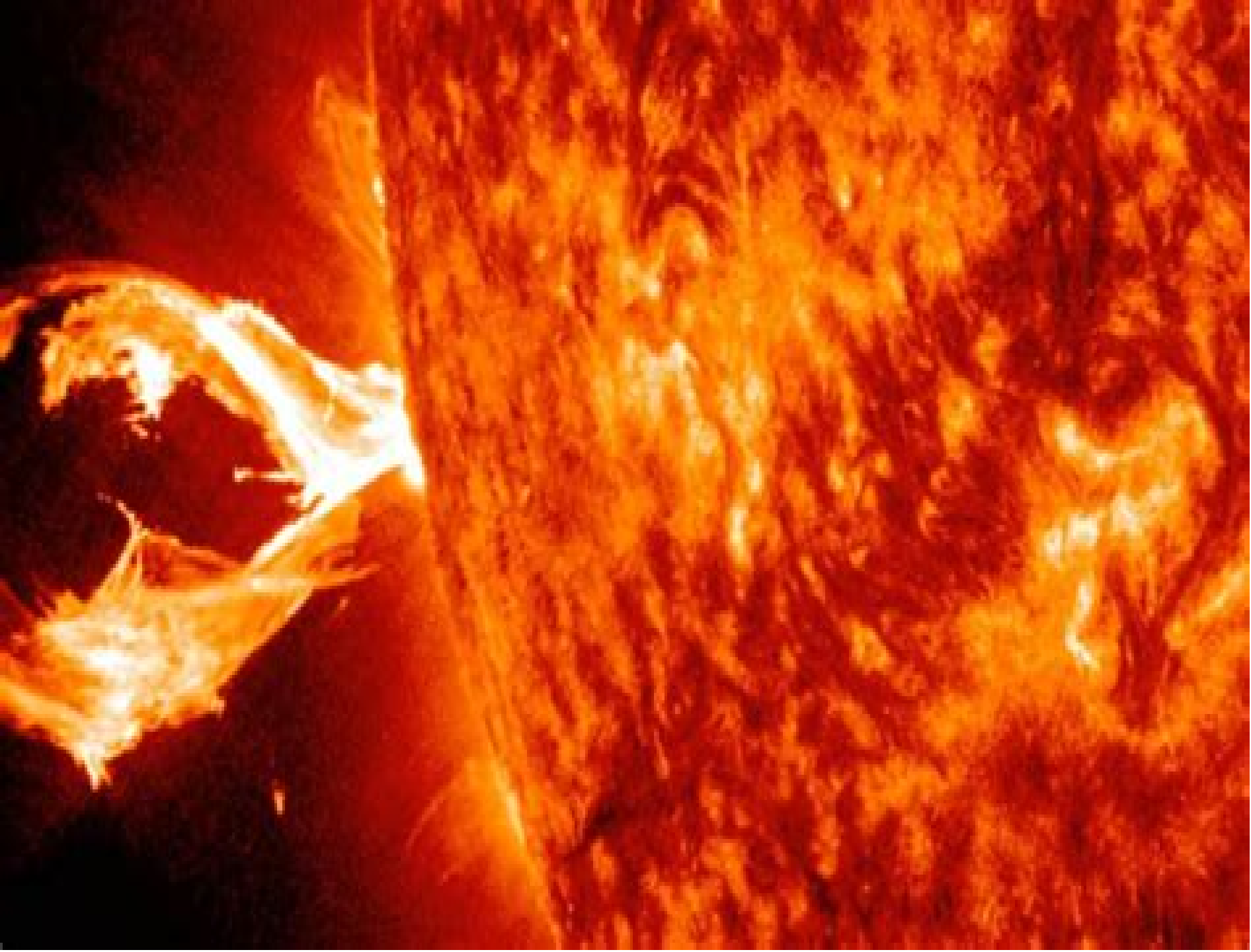




Solar Flare

Yuhong Fan, George Fisher



Solar Flare:

Solar Flares Zdenek Svestka, 2012-12-06 This book is the first part of the originally planned publication by Z Svestka and L D de Feiter Solar High Energy Photon and Particle Emission The second part with the original title was to be published by de Feiter in about one year from now However to the deep sorrow of all of us Dr de Feiter died suddenly and unexpectedly when the present book was in print Thus unfortunately de Feiter's second part may not appear Due to the fact that the originally planned publication was divided into two parts the present book is mainly descriptive and concerned with the flare morphology It was expected that theoretical interpretations would be extensively developed in the second part prepared by de Feiter In particular this refers to the theoretical back grounds of radio emissions particle acceleration and particle propagation in space Only in Chapter II concerning the low temperature flare do we go deeper into the theoretical interpretations anticipating that de Feiter would have been concerned mainly with the high energy physics Still the book includes discussions on all important aspects of flares and thus can present the reader with a complete picture of the complex flare phenomenon It is clear that many observed data on flares can be interpreted in different ways

Solar Flare Magnetic Fields and Plasmas Yuhong Fan, George Fisher, 2012-03-28 This volume is devoted to the dynamics and diagnostics of solar magnetic fields and plasmas in the Sun's atmosphere Five broad areas of current research in Solar Physics are presented 1 New techniques for incorporating radiation transfer effects into three dimensional magnetohydrodynamic models of the solar interior and atmosphere 2 The connection between observed radiation processes occurring during flares and the underlying flare energy release and transport mechanisms 3 The global balance of forces and momenta that occur during flares 4 The data analysis and theoretical tools needed to understand and assimilate vector magnetogram observations and 5 Connecting flare and CME phenomena to the topological properties of the magnetic field in the Solar Atmosphere The role of the Sun's magnetic field is a major emphasis of this book which was inspired by a workshop honoring Richard C Dick Canfield Dick has been making profound contributions to these areas of research over a long and productive scientific career Many of the articles in this topical issue were first presented as talks during this workshop and represent substantial original work The workshop was held 9-11 August 2010 at the Center Green campus of the National Center for Atmospheric Research NCAR in Boulder Colorado This volume is aimed at researchers and graduate students active in solar physics solar terrestrial physics and magneto hydrodynamics Previously published in Solar Physics journal Vol 277 1 2012

Hard X-Ray Imaging of Solar Flares Michele Piana, A. Gordon Emslie, Anna Maria Massone, Brian R. Dennis, 2022-01-01 The idea for this text emerged over several years as the authors participated in research projects related to analysis of data from NASA's RHESSI Small Explorer mission The data produced over the operational lifetime of this mission inspired many investigations related to a specific science question the when where and how of electron acceleration during solar flares in the stressed magnetic environment of the active Sun A vital key to unlocking this science problem is the

ability to produce high quality images of hard X rays produced by bremsstrahlung radiation from electrons accelerated during a solar flare The only practical way to do this within the technological and budgetary limitations of the RHESSI era was to opt for indirect modalities in which imaging information is encoded as a set of two dimensional spatial Fourier components Radio astronomers had employed Fourier imaging for many years However differently than for radio astronomy X ray images produced by RHESSI had to be constructed from a very limited number of sparsely distributed and very noisy Fourier components Further Fourier imaging is hardly intuitive and extensive validation of the methods was necessary to ensure that they produced images with sufficient accuracy and fidelity for scientific applications This book summarizes the results of this development of imaging techniques specifically designed for this form of data It covers a set of published works that span over two decades during which various imaging methods were introduced validated and applied to observations Also considering that a new Fourier based telescope STIX is now entering its nominal phase on board the ESA Solar Orbiter it became more and more apparent to the authors that it would be a good idea to put together a compendium of these imaging methods and their applications Hence the book you are now reading *Constructing Solar Flare Events by Computer* Hans J. E. Fischer, 1969 A computer program has been developed which processes and groups individual observations of solar flares which describe the same feature From each group of observations a consolidated description of the flare event is produced The report details the methods by which these groups are formed and the subsequent consolidated descriptions achieved Author **Solar Flare Occurrence as a Function of Sunspot Size** Isadore Enger, 1966 An objective technique has been developed for assessing the probability of occurrence of solar flares in a particular sunspot group given only the sizes of the sunspot group on two consecutive days The sunspot data cover the 10 year period 1955 to 1964 and have been published in the Greenwich Photoheliographic Results The flare data were measured at Sacramento Peak New Mexico The flare and spot data were combined by a simple objective technique which depends only on the size of the spot group and the relative positions of the spot group and flare These results show the very marked dependence of flare occurrence on the size of the spot group and a significant but very much smaller dependence on the change in group size from one day to the next Author **Nuclear Science Abstracts**, 1976-06 **Flare Observations in the IRIS Era: What Have We Learned, and What's Next?** Vanessa Polito, Bart de Pontieu, Adam Kowalski, Viggo Hansteen, 2023-12-27 Since its launch in 2013 IRIS has observed more than 10 X class over 100 M class and more than 600 C class flares at unprecedented spatial and temporal resolution Thanks to the rich diagnostics that cover the physical conditions of the solar atmosphere from the photosphere to the hottest parts of the flaring corona IRIS observations have provided exciting new results and constraints on flare heating models significantly expanding our knowledge of how flares are triggered and how the non thermal energy is released propagates downward from the corona and is deposited in the low atmosphere At the same time the new discoveries provided by IRIS have raised new unresolved questions and new

challenges for theoretical models For instance current hydrodynamic models still cannot fully explain many features observed by IRIS during both the impulsive and gradual phases such as the dynamics of the evaporative condensation flows the large line broadenings and the puzzling complex and broad chromospheric lines In addition important questions remain regarding the details of the energy propagation and dissipation in flares the importance of Alfvén waves vs electron beam and thermal conduction heating and the effects from large scale reconfiguration of the magnetic field during flares

Scientific and Technical Aerospace Reports ,1990 **Energy Research Abstracts** ,1990 **Physical Processes in Solar Flares**

B.V. Somov,1991-12-31 Solar flares are very complex electromagnetic phenomena of a cataclysmic nature Particles are accelerated to very high velocities and a variety of physical processes happen inside and outside flares These processes can be studied by a large number of techniques from Earth and from space The aim is to discover the physics behind solar flares This goal is complicated because information about the flare mechanism can be obtained only in an indirect way by studying the secondary effects This book provides three stages in the solution of the solar flare problem Chapter one describes the connection between observational data and theoretical concepts where it is stressed that next to investigating flares the related non stationary large scale phenomena must be studied as well The second chapter deals with secondary physical processes in particular the study of high temperature plasma dynamics during impulsive heating The last chapter presents a model built on the knowledge of the two previous chapters and it constructs a theory of non neutral turbulent current sheets The author believes that this model will help to solve the problem of solar flares For solar physicists plasma physicists high energy particle physicists

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